



National Aeronautics and Space  
Administration  
Jet Propulsion Laboratory  
California Institute of Technology

# Deep Space Optical Communications Visions, Trends and Prospects

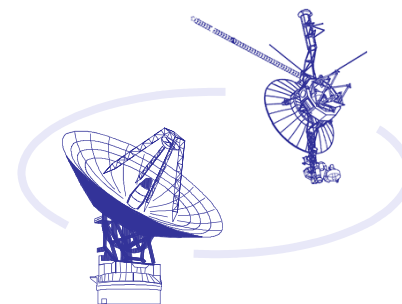
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# Introduction

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- Background
- Drivers: Planetary Science; Astrophysics; Human Exploration
- Data Rates & Data Rate Trends
- Spectrum
- Optical & RF Comparisons
- SCaN's *Optical Communications Roadmap*
- Early Steps
- Summary



# Background

- Interest in optical communications has certainly grown since the 1980s
  - Its potential is not hard to recognize
- However, implementation, and even demonstration, has been a “hard sell”
  - Do customers really need it? Yet?
  - They’d rather exploit the remaining RF “head room”
- Customer’s demand side
  - Moving to a different class of exploratory missions
  - Rapid pace of science instrument development
  - New discoveries arise from higher spatial, temporal and spectral resolution
- Technology developer’s side
  - Technology may be leading the customer needs
  - But something new always entails more risk
  - Roles of partnership vis-à-vis competition
- Policymaker’s side
  - Can encourage or discourage customer demand
  - Outreach demand may trump science demand
- Von Braun analogy re: the state of rocketry in the early-1950s
  - Reference: "A Plea for a Coordinated Space Program," article in *The Complete Book of Outer Space*, published 1953 by the Gnome Press.



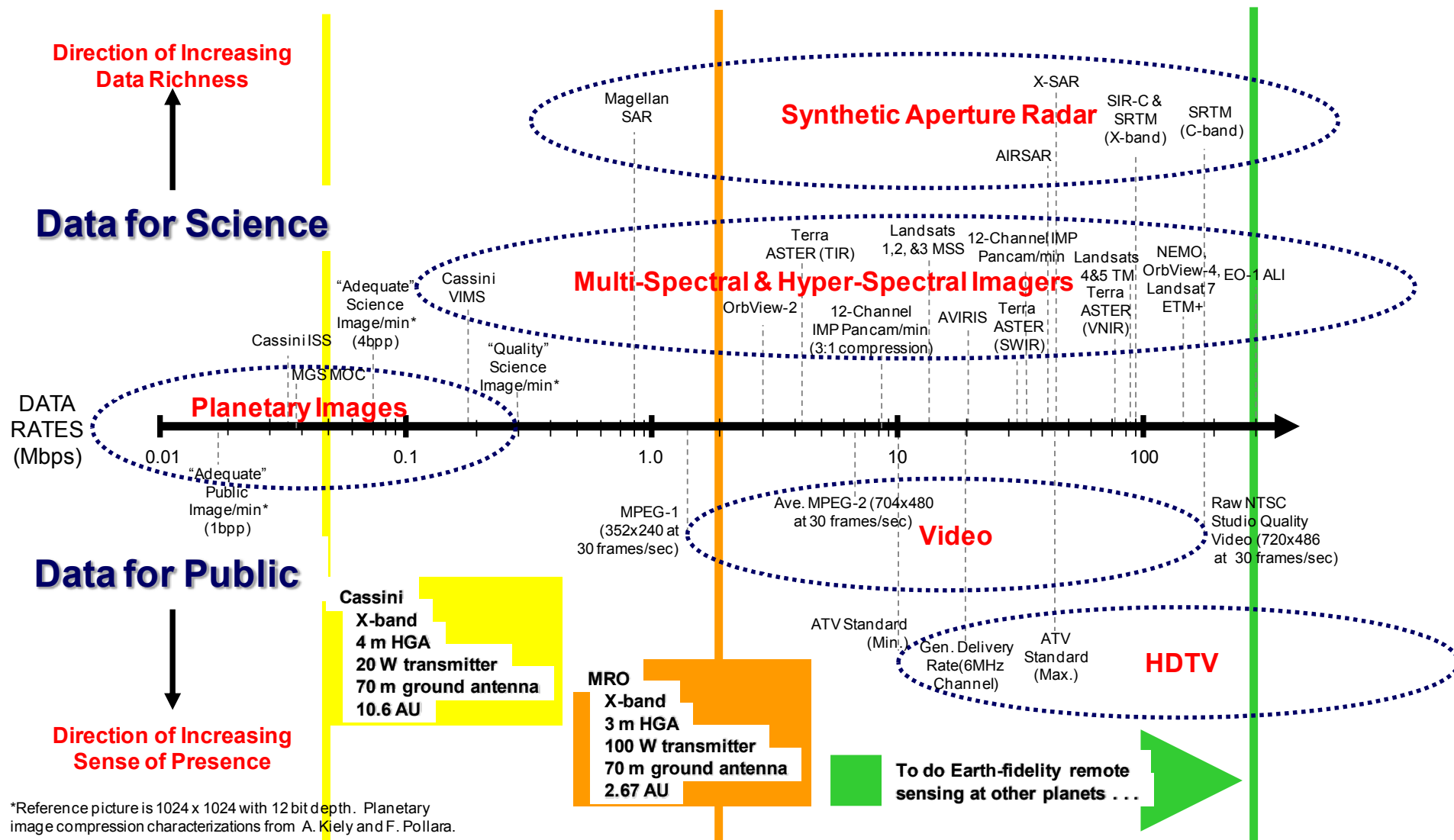
# Drivers: Planetary Science

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- 1980s – early-1990s: Tail end of the initial reconnaissance of the solar system
  - Most demanding deep space missions typically had D/L data rates ~10s to 100s of kbps
  - Sufficient to return first images of other planets
  - No threat to consuming allocated spectrum
  - Links were challenging but engineers still envisioned numerous RF improvements
    - Higher RF frequencies
    - Better FEC coding
    - Lower receiver noise temperature
    - Larger receiving area
    - Greater EIRP on spacecraft
- Mid-1990s - Today: Re-examine planetary targets in more detail
  - Preliminary reconnaissance of the solar system has essentially been completed
    - All planets had been visited at least once (*Note: Pluto got demoted!*)
  - Current deep space missions need D/L data rates ~100s of kbps to 10s of Mbps, i.e., more than an order of magnitude increase
    - Images: higher resolution and/or multi-spectral
    - SAR observations
    - Near R/T video
  - Remote sensing of other planets, at the same fidelity done at Earth today, requires an increase of more than three orders of magnitude



# Required Data Rates as a Function of Data Type



# Other Drivers

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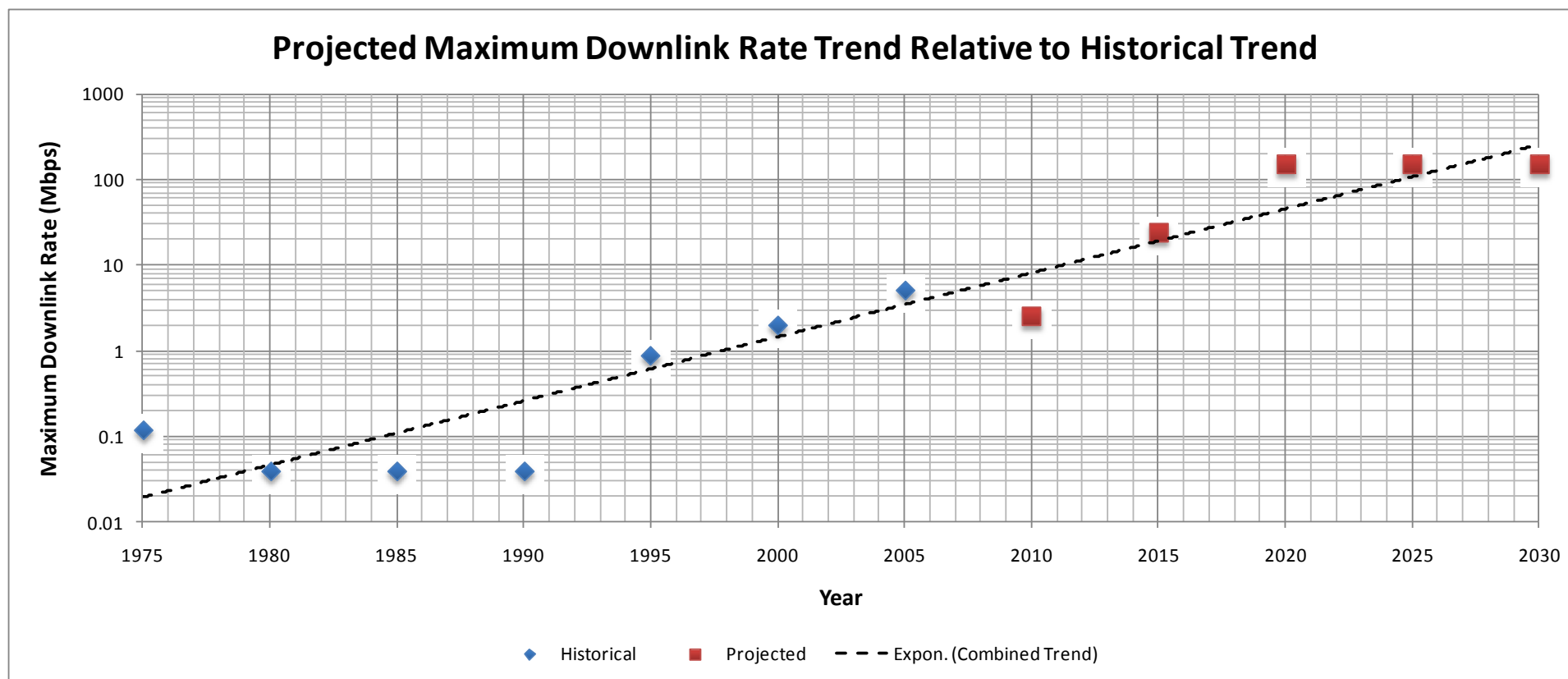
## Astrophysics:

- 1990s – early-2000s: NASA's Great Observatories
  - Spitzer (IR); Hubble (Visual); Chandra (X-ray); Compton ( $\gamma$ -ray)
  - Typical D/L data rates: 0.5 Mbps to 2.0 Mbps
- 2010 - 2020: Greater (?) Observatories
  - JWST D/L data rates: 25 Mbps, i.e., more than an order of magnitude increase
- 2020s - ??? : Greatest (?) Observatories
  - Concepts for dark energy investigation D/L data rates: 150 Mbps, i.e.,  $\sim 2$  orders of magnitude beyond the Great Observatories

## Human Exploration:

- Late-1960s – early-1970s: Apollo era
  - S-band D/L data rates: 50 kbps
- 1980s - 2020: Space Shuttle / ISS era
  - Ku-band D/L data rates: 50 Mbps
- 2020s - ??? : Lunar return / Near-Earth Objects / Mars expedition
  - Anticipated Ka-band D/L data rates: 150 Mbps

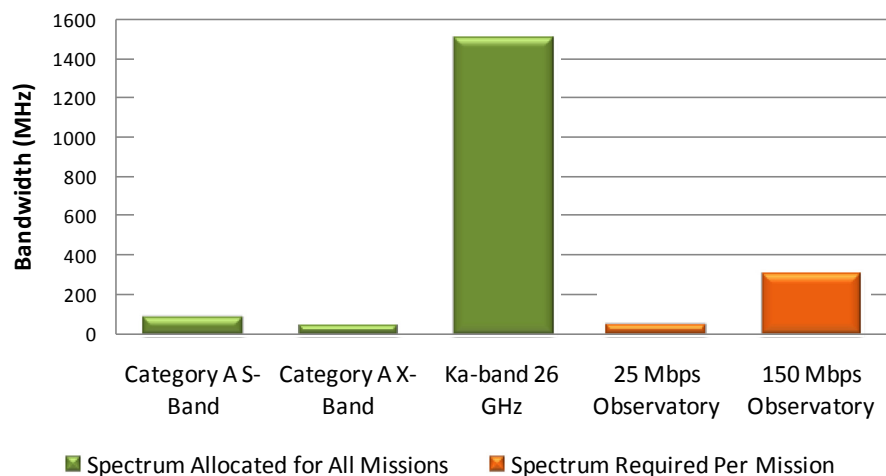
# Historical and Projected Downlink Rate Trend



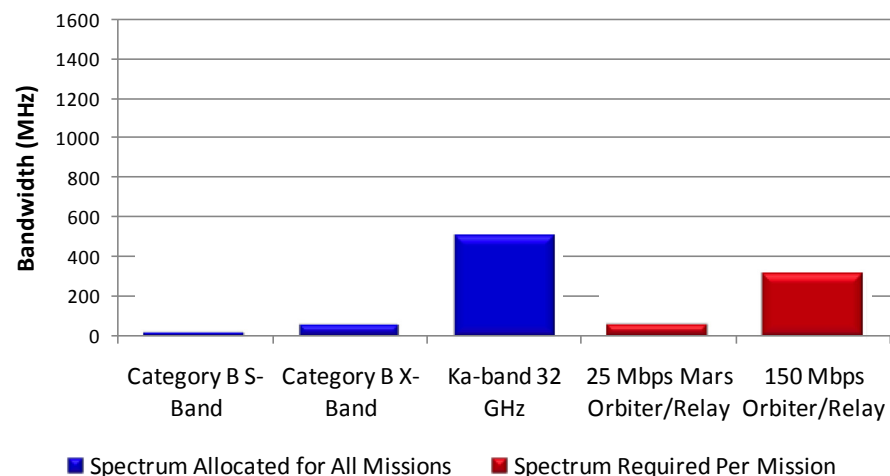


## Category A and B Spectrum Allocations Relative to High-Rate Mission Bandwidth Requirements

**Spectrum Implications: Future Missions  
within 2 Million km of Earth**



**Spectrum Implications: Future Missions  
beyond 2 Million km from Earth**





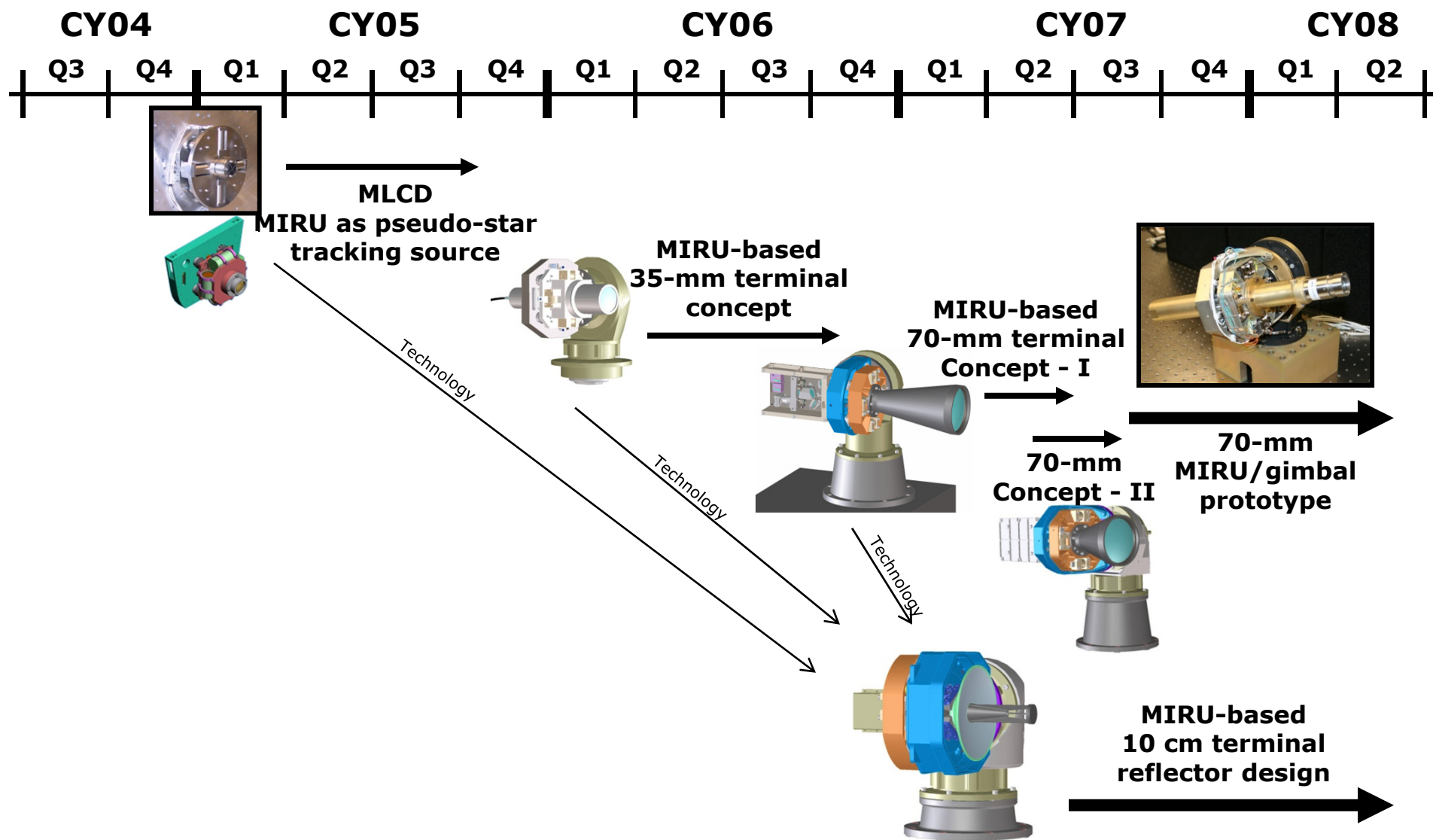
# Future Downlink Possibilities at RF and Optical

	Data Rate Today		Data Rate ~2020		Data Rate ~2030		
Spacecraft Capabilities	3m Antenna X-Band 100 W Xmitter		3m Antenna Ka-Band 180 W Xmitter		5m Antenna Ka-band 200 W Xmitter	1m Optical 1550 nm 50 W Xmitter	
DSN Antennas	1 x 34m	3 x 34m	1 x 34m	Equiv to 3 x 34m	1 x 34m	Equiv to 7 x 34m	10m Optical
Mars (0.6 AU)	20 Mbps	60 Mbps	400 Mbps	*1.2 Gbps	*1.3 Gbps	*9.3 Gbps	5.5 Gbps
Mars (2.6 AU)	1 Mbps	3 Mbps	21 Mbps	64 Mbps	71 Mbps	*500 Mbps	300 Mbps
Jupiter	250 Kbps	750 Kbps	5 Mbps	15 Mbps	16 Mbps	115 Mbps	70 Mbps
Saturn	71 Kbps	213 Kbps	1.4 Mbps	4 Mbps	4.7 Mbps	33 Mbps	19 Mbps
Neptune	8 Kbps	24 Kbps	160 Kbps	470 Kbps	520 Kbps	3.7 Mbps	2.2 Mbps

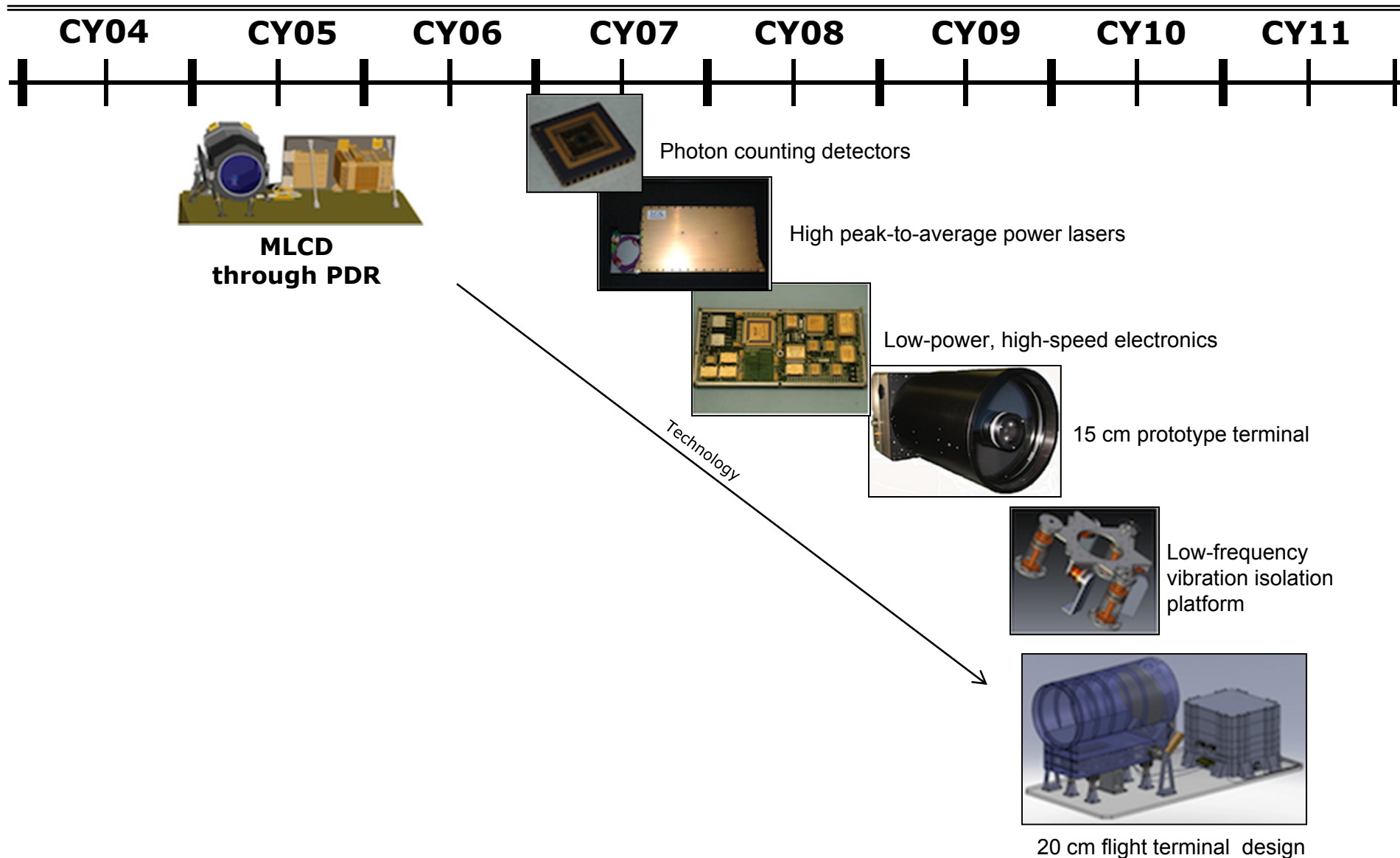
- \* Reference spacecraft is MRO-class (power and antenna), Rate 1/6 Turbo Coding, 3 dB margin, 90% weather, and 20° DSN antenna elevation
- \*\* Performance will likely be 2 to three times lower due to need for bandwidth-efficient modulation to remain in allocated spectrum



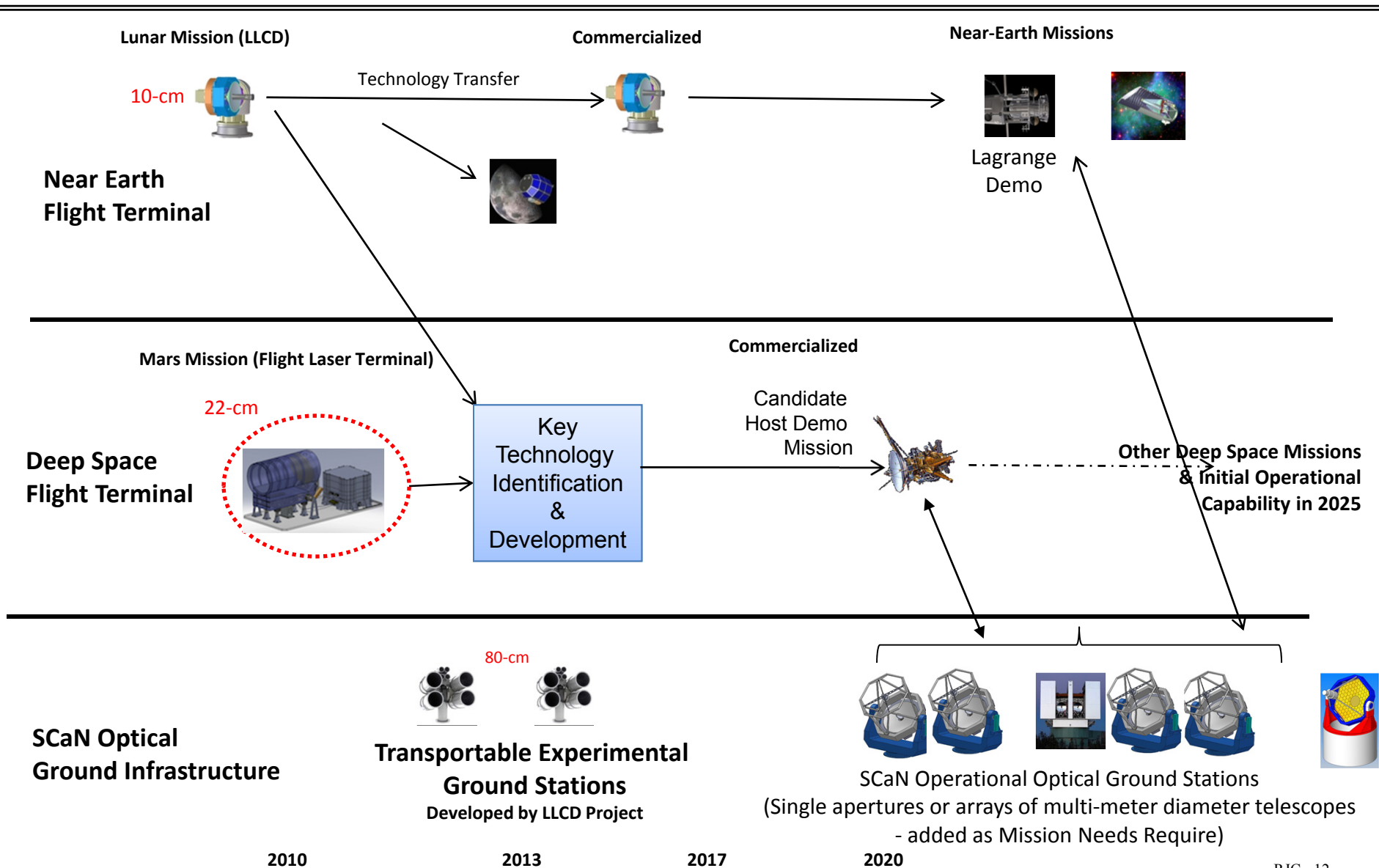
# SCaN Optical Program Background



# SCaN Deep Space Optical Program Background



# NASA Strategy for Optical Communications Development



# SCaN's Top-Level Demonstration Objectives

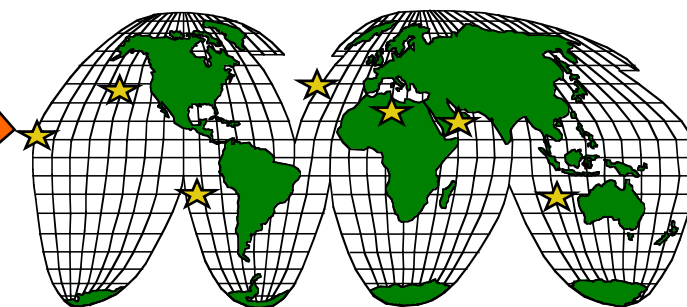
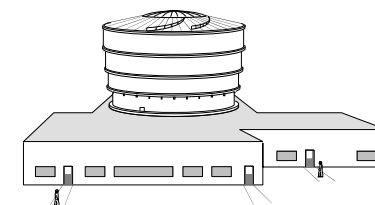
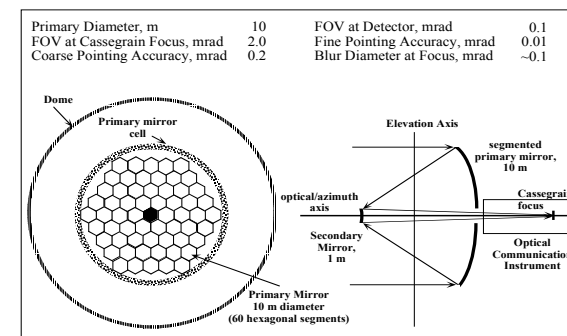


Objective	LLCD	Lunar Lander	L1	L2	Mars
High Data Rate (10X RF)					
Pointing, Acquisition & Tracking for Lunar/L1/L2					N/A
Pointing, Acquisition & Tracking for Deep Space					
Day Time Reception at Ground Terminals					
Low SEP Downlink Acquisition					
Low SPE Uplink Beacon Acquisition					
Lifetime in space					
Weather & Ground Station Handover					



# 1993 Ground-Based Antenna Technology Study (GBATS)

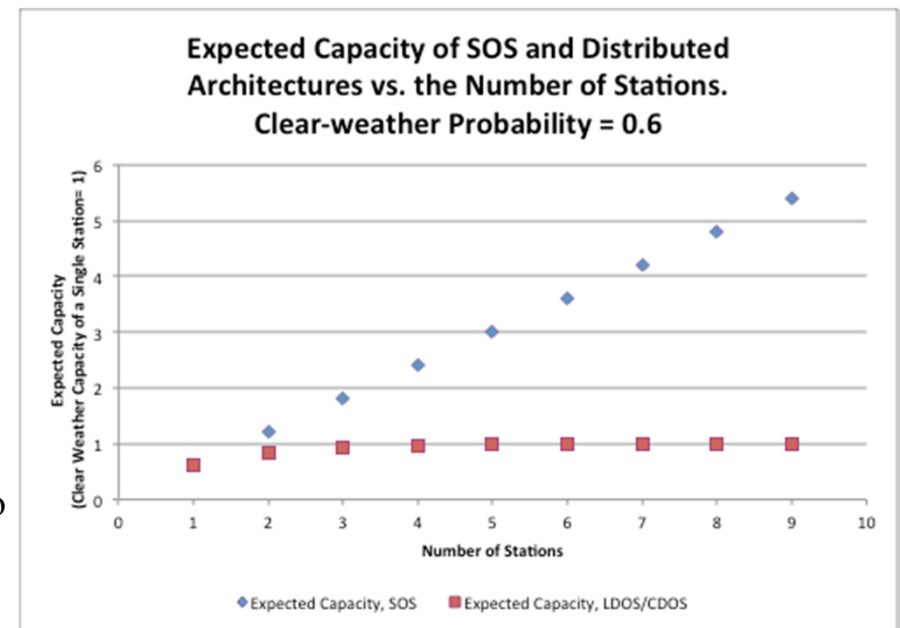
- Spatially-diverse network of optical ground stations
- 10m diameter, segmented aperture photon buckets
  - Also included 1m uplink telescope stations
- Station and network infrastructure
- Options considered
  - Clustered Optical Subnet (COS): 3 longitude regional subnets; 3 spatially-diverse stations each
  - Linearly Dispersed Optical Subnet (LDOS): N-stations around Earth
- Study recommendations:
  - 6 to 8 station LDOS judged as best
  - Best 24-hour availability at lowest cost



7-Station LDOS

## Early Step: Single Optical Site

- LDOS (& COS) were close to ideal ground network architectures
  - High availability enables traditional ConOps
  - But high cost and geopolitical issues remain as barriers
- Single Optical Site (SOS) proposes a ConOps paradigm shift
  - Remove (at least temporarily) requirement for high availability optical D/L
  - Replace with top-level requirement for maximization of science data return
  - Utilize optical link for high-B/W, high value but low temporal priority science data
  - Utilize RF links for routine TT&C, thumbnail science, critical event and emergency support
  - Note: SOS can always be upgraded to an LDOS or COS in the future
- SOS has some unanticipated characteristics
  - Aggregating photon collection capability at a single site is much more efficient than dispersing it to increase availability
    - 3X improvement in science data return is typical in comparison to a 5-station LDOS/CDOS architecture
  - However, provisions must be made for retransmissions with on-board solid state recorders
    - State-of-the-art in recorders is adequate – with the possible exception of Jovian radiation cases
  - Links from a Mars lander may be problematic due to nearly synchronous rotation rates of Mars and Earth







## Early Step: RF-Optical Hybrid

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- Modify DSN 34m X/Ka-band (8/32 GHz) antennas for reception of optical signals
  - Preliminary results show promise that dual RF-optical may be possible on the same ground terminal
  - Operational and cost benefits can result from dual use of the same aperture
  - The *utmost* in network integration - a current priority for the SCA Office
  - Antennas being considered have: robust backup structures; large collecting areas; and millidegree pointing - all of which support optical communications
- Candidate design concept: polish / coat the inner 26m-diameter aluminum panels of a 34m antenna to a high degree of reflectivity
  - Though panels are optically smooth, they will still have underlying surface imperfections
    - Will generate large (several cm) spots at the Cassegrain focus corresponding to a FOV of hundreds of  $\mu\text{rad}$
  - Large-area photon-counting-detector arrays convert the optical fields to photon counts for downstream digital processing.
  - A solar energy filter over the main reflector protects the antenna from sunlight and the panels from dust.
- Candidate design concept: replace some panels with optical reflectors
  - Optical surfaces (either monolithic or arrayed) have aperture equivalent to a 10m terminal
  - Relies on high-quality glass mirrors that replace a fraction of the aluminum panels of the antenna
  - Achieves a much smaller optical FOV while still maintaining adequate RF performance
    - Mirrors will generate much smaller spots, typically limited by turbulence to  $\sim 50 \mu\text{rad}$  FOV.
  - Use of spherical mirrors, given large overall antenna focal length, reduces implementation cost.
  - As in the other concept, a solar energy rejection filter provides protection from heat and dust.



## Two Other Relevant Factors: One a minus; the other a plus

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- Data, more data and even more data!
  - Data generation by missions, as well as by ground based investigations, continues to grow exponentially
  - Are we already awash in too much data?
  - Can we process all these data and mine them for useful knowledge?
  - Does it make sense to archive Tbytes of data that no one will ever examine?
  - If future mission operations concepts pre-select a small subset of collected data for downlink, that could reduce support for optical communications
- Commercial industry and spectrum
  - Demand for microwave spectrum by commercial entities seems to be insatiable
  - If deep space RF spectrum allocations become threatened, that could increase support for optical communications



# Meet the New Competitors for RF Bandwidth!

- UAV use is proliferating; their ISR data transfer needs are driving a migration to X- and Ka-band.
- More “hot spots” around the globe are driving up VSAT requirements and associated bandwidth demand.
- Military and commercial information devices are growing smaller and more ubiquitous, with some of the supporting links driving up bandwidth demand. Extended Ku-band already being eyed. Ka-band next.
- Government use of commercial space assets growing to ensure network resilience.
- Commercial satellite providers increasingly offering hosted payloads to government users





# Summary

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- Justification for deep space optical communications is abundantly clear – at least to us!
  - Ever-growing mission requirements for data rates
  - Spectrum needed to accommodate such rates
- However, RF communications still have some potential for growth
  - Missions will prefer to exploit this rather than make the riskier leap to optical links
- Ultimately, there will be no alternative to optical communications
  - ‘Ultimately’ can be a long time; need to make things happen sooner
- Two strategies inherent in the *SCaN Optical Communications Roadmap*
  - Continue to invest in technologies that will improve performance, operability, risk and cost
  - Validate these technologies via demonstrations in the relevant environments
- Explore novel ConOps that may lower the cost of optical systems – in essence reducing the barriers to entry
  - This might not provide the ideal capability at the start
  - It can provide a foundation upon which to grow for the future
- Von Braun analogy
  - Although prospects may appear bleak at times, the “window of opportunity” will open
  - You have to be ready when it does!